

Calibrate Shear Multiplicative Bias with CMB Lensing

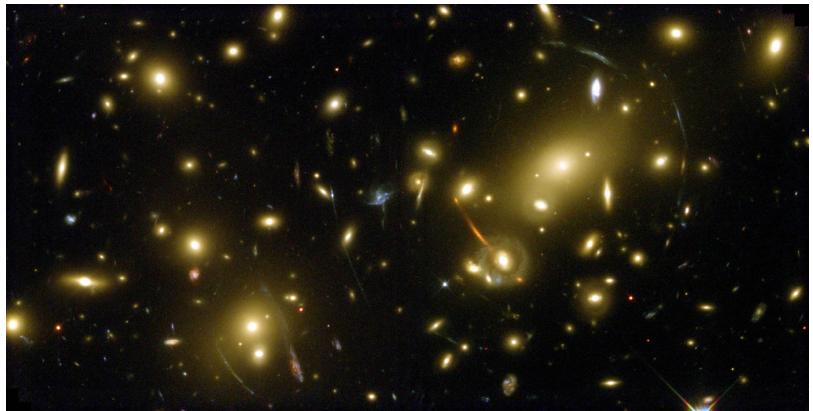
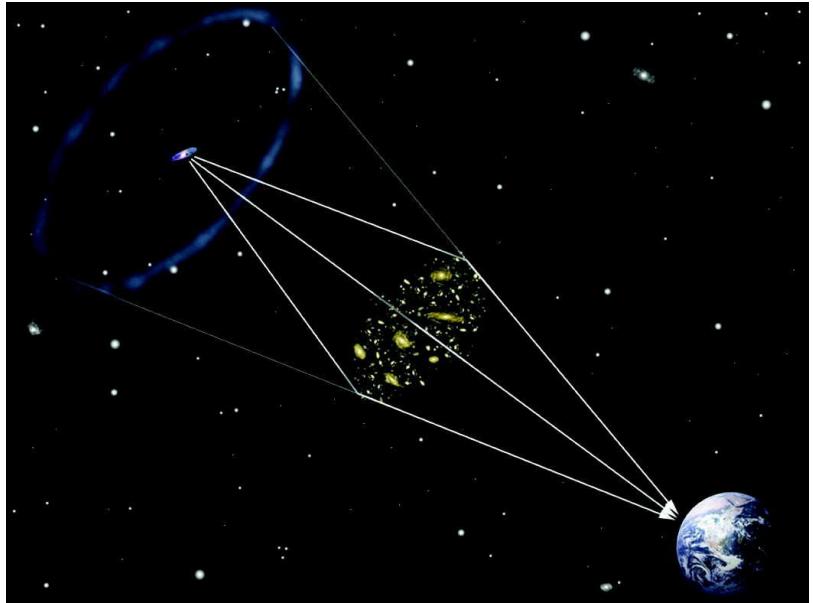
Jia Liu (Columbia Univ.)

CCS @ BNL/Stony Brook, May 23rd 2016

1503.06214: JL & Hill

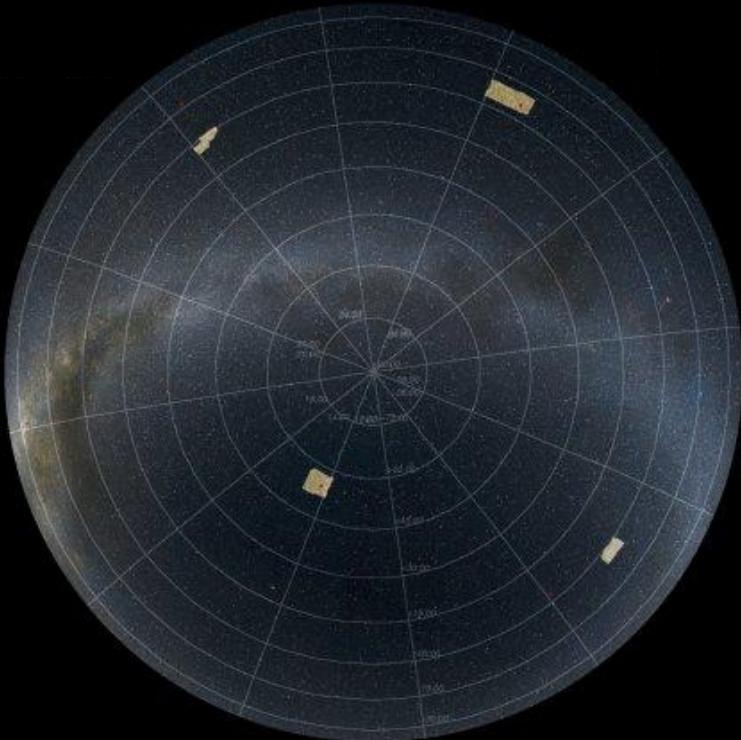
1601.05720: JL, Ortiz-Vazquez, & Hill

Weak Lensing



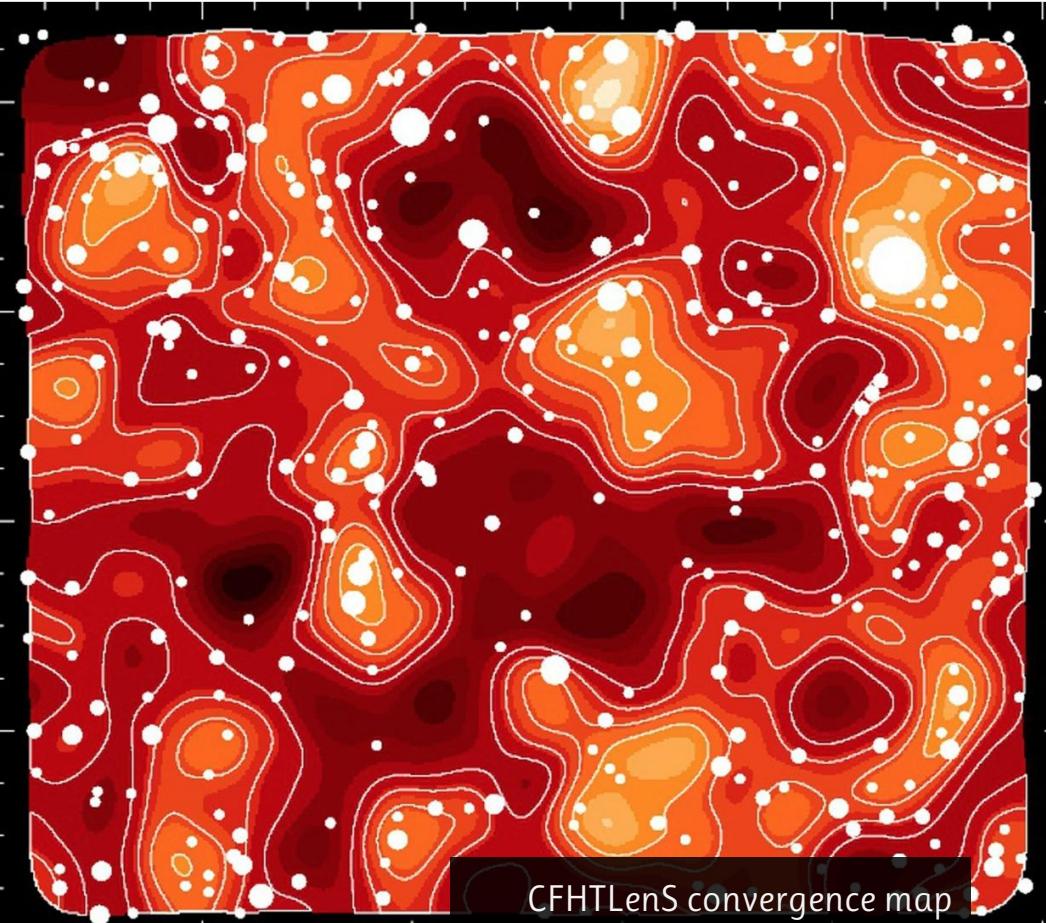
Credit: wikipedia (top), CFHTLenS (bottom)

154 deg² Canada-France-Hawaii Telescope Lensing Survey (CFHTLenS)



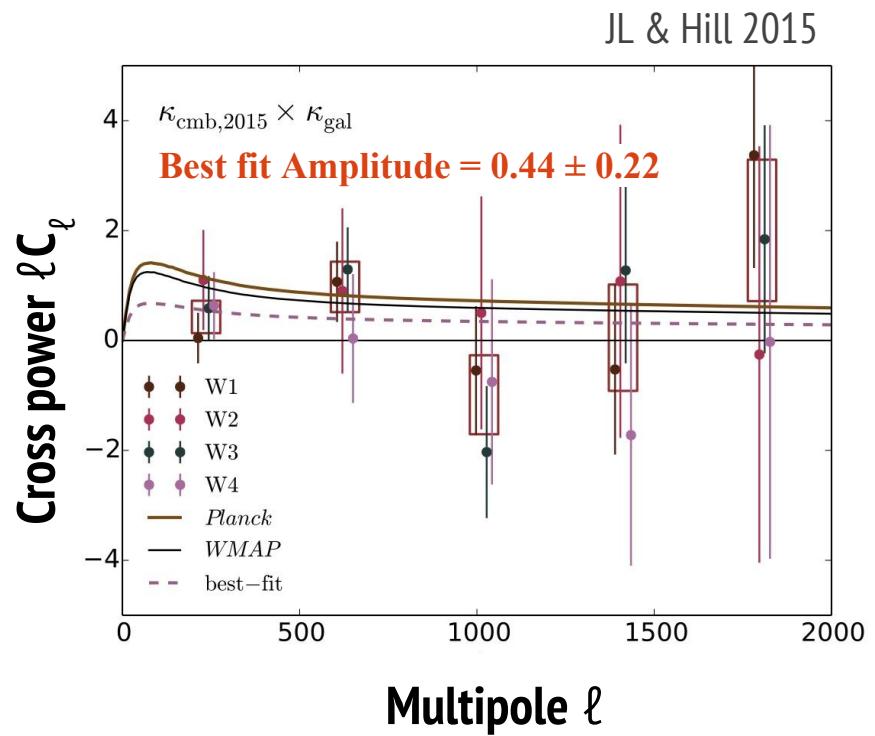
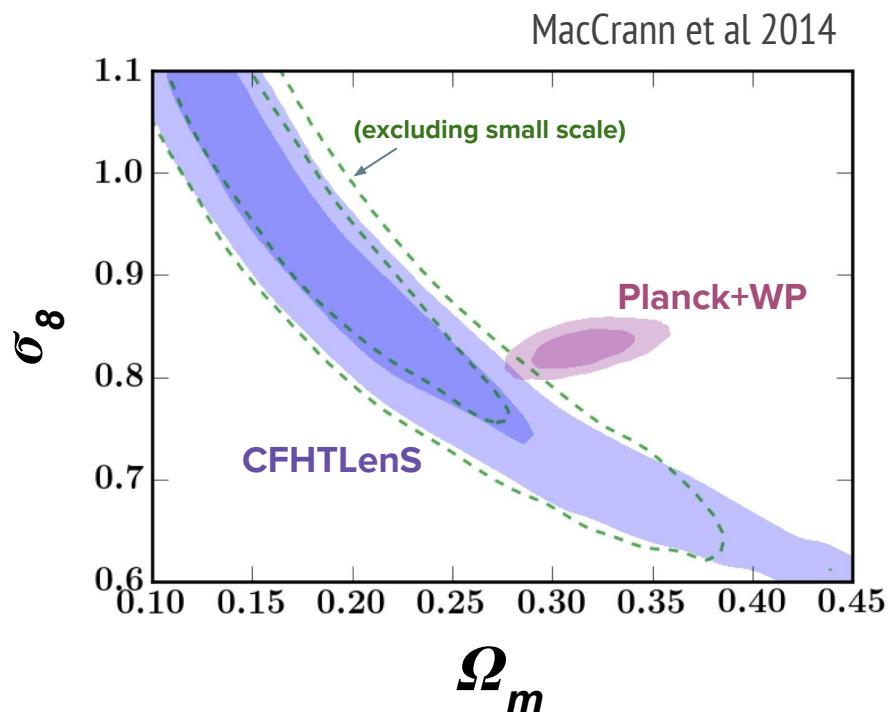
CFHTLS fields across the northern sky

- Deep
- Wide

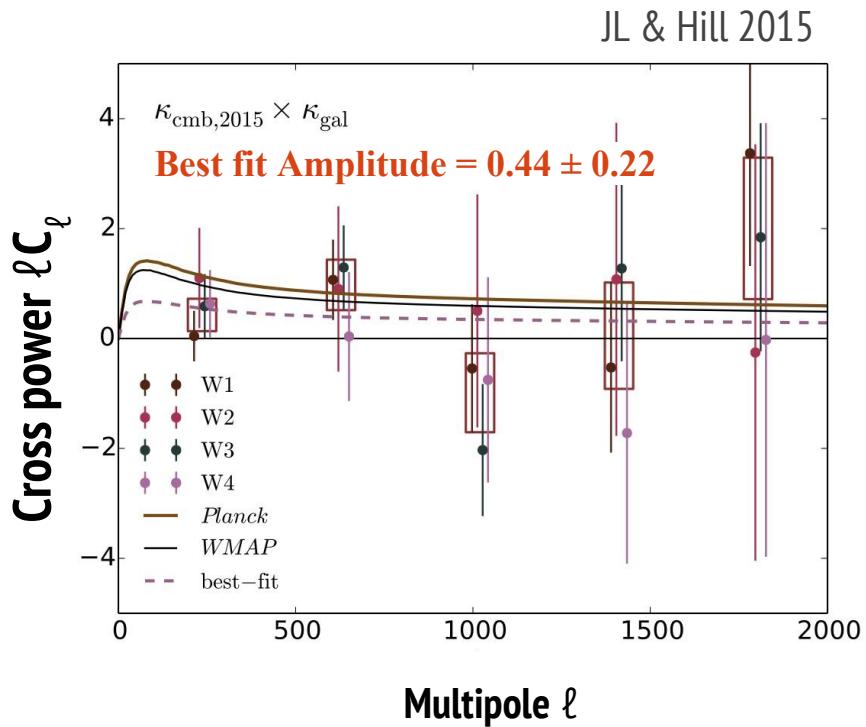


CFHTLenS convergence map
Van Waerbeke et al 2013

2σ Discrepancies between CMB Temp. & weak lensing



Sources of the 50% disagreement?



Photometric Redshift (10%)

Intrinsic Alignments (10-15%)

Masking of tSZ Clusters (5-10%)

Multiplicative Bias (?)

New Physics (maybe not..yet)

The Multiplicative Bias

Model Bias

Noise Bias

Model & Noise
Bias Coupling

The Multiplicative Bias

Model Bias

e.g. Bernstein 2010
Voigt et al 2012

Noise Bias

Model & Noise
Bias Coupling



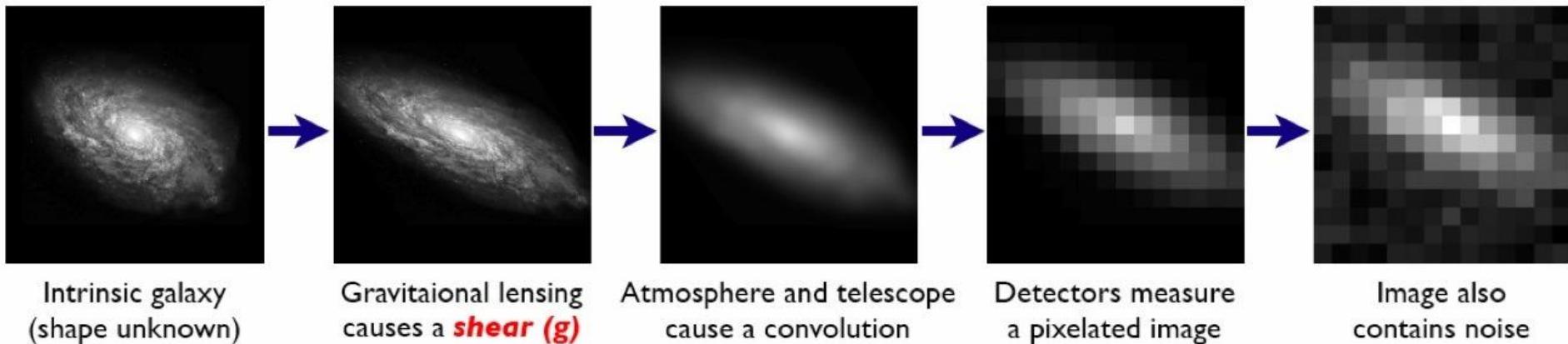
The Multiplicative Bias

Model Bias

Noise Bias

Model & Noise
Bias Coupling

e.g. Melchior & Viola 2012
Okura & Futamase 2013



Credit: Catherine Heymans

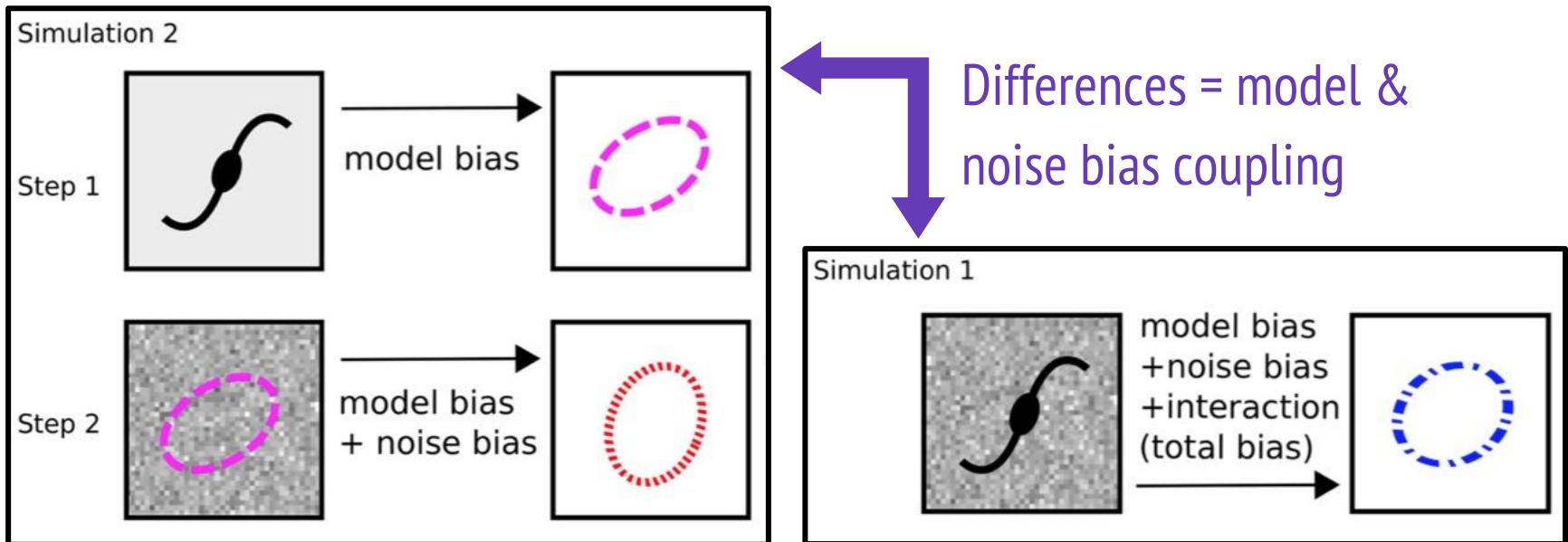
The Multiplicative Bias

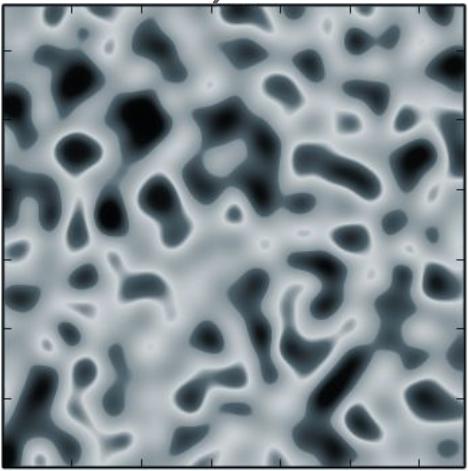
Model Bias

Noise Bias

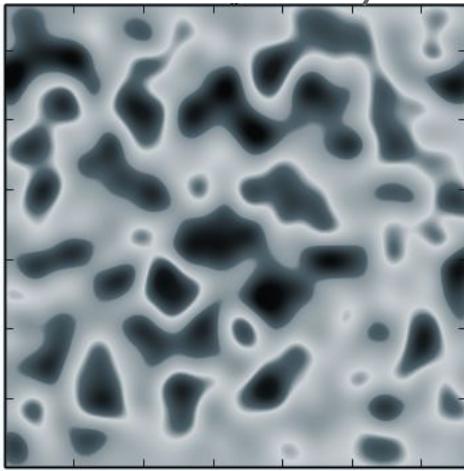
Model & Noise
Bias Coupling

Kacprzak et al 2014

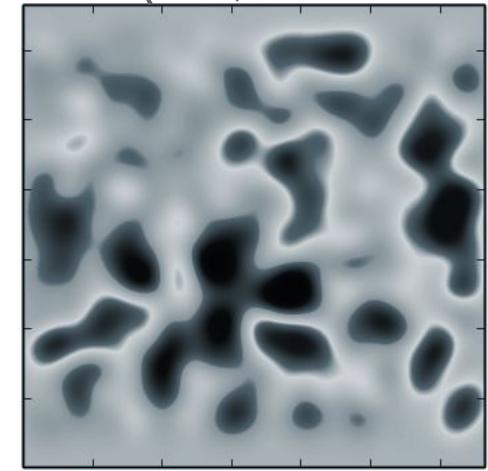




Planck
CMB lensing



CFHTLenS
galaxy lensing



CFHTLenS
galaxy density

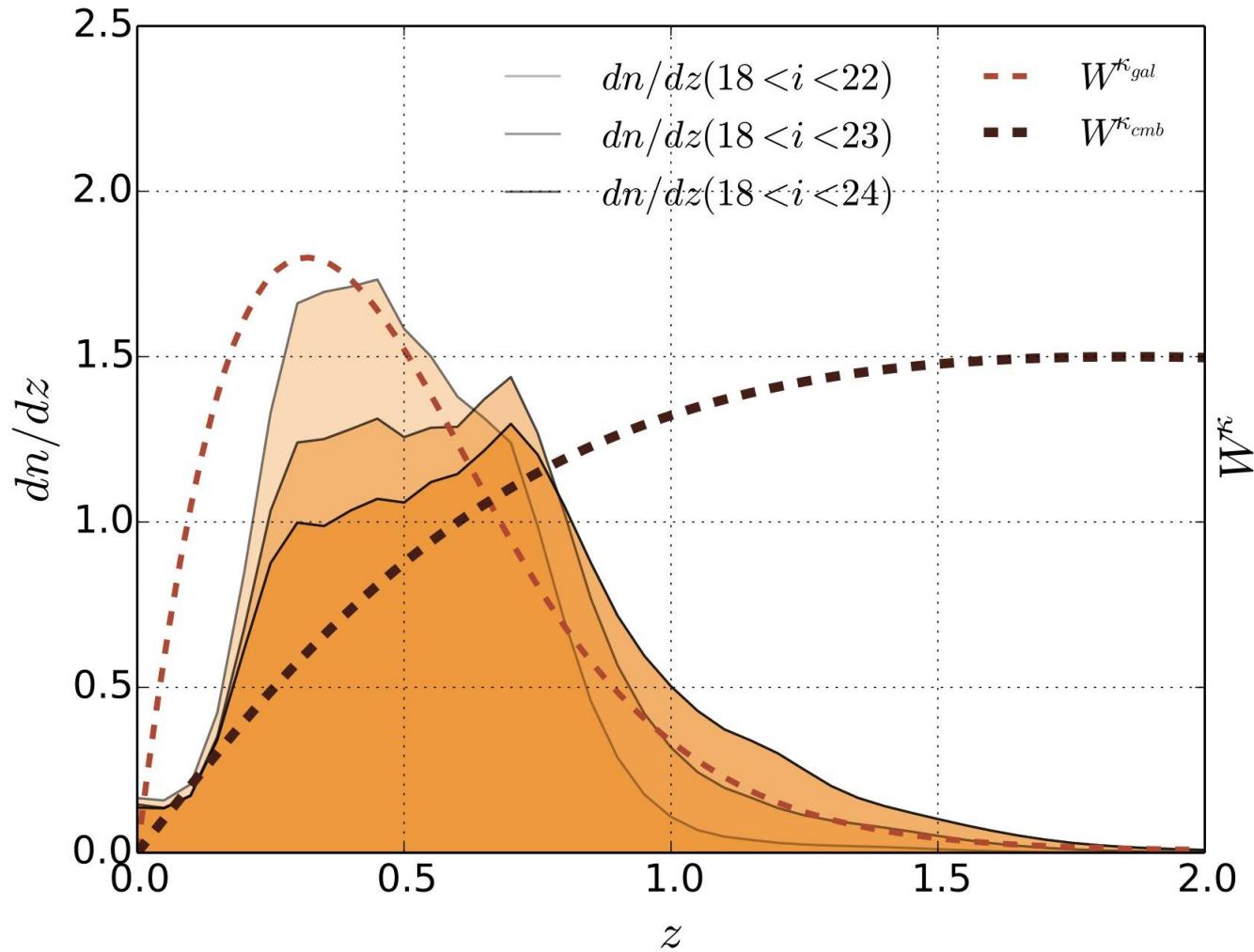
A $C_\ell^{\kappa_{\text{CMB}} \Sigma} = \frac{3}{2} \Omega_m H_0^2 \int d\eta b_\ell(\eta) W_f(\eta) \frac{g_{\text{CMB}}(\eta)}{a(\eta)} P\left(\frac{\ell}{d_A}, \eta\right),$

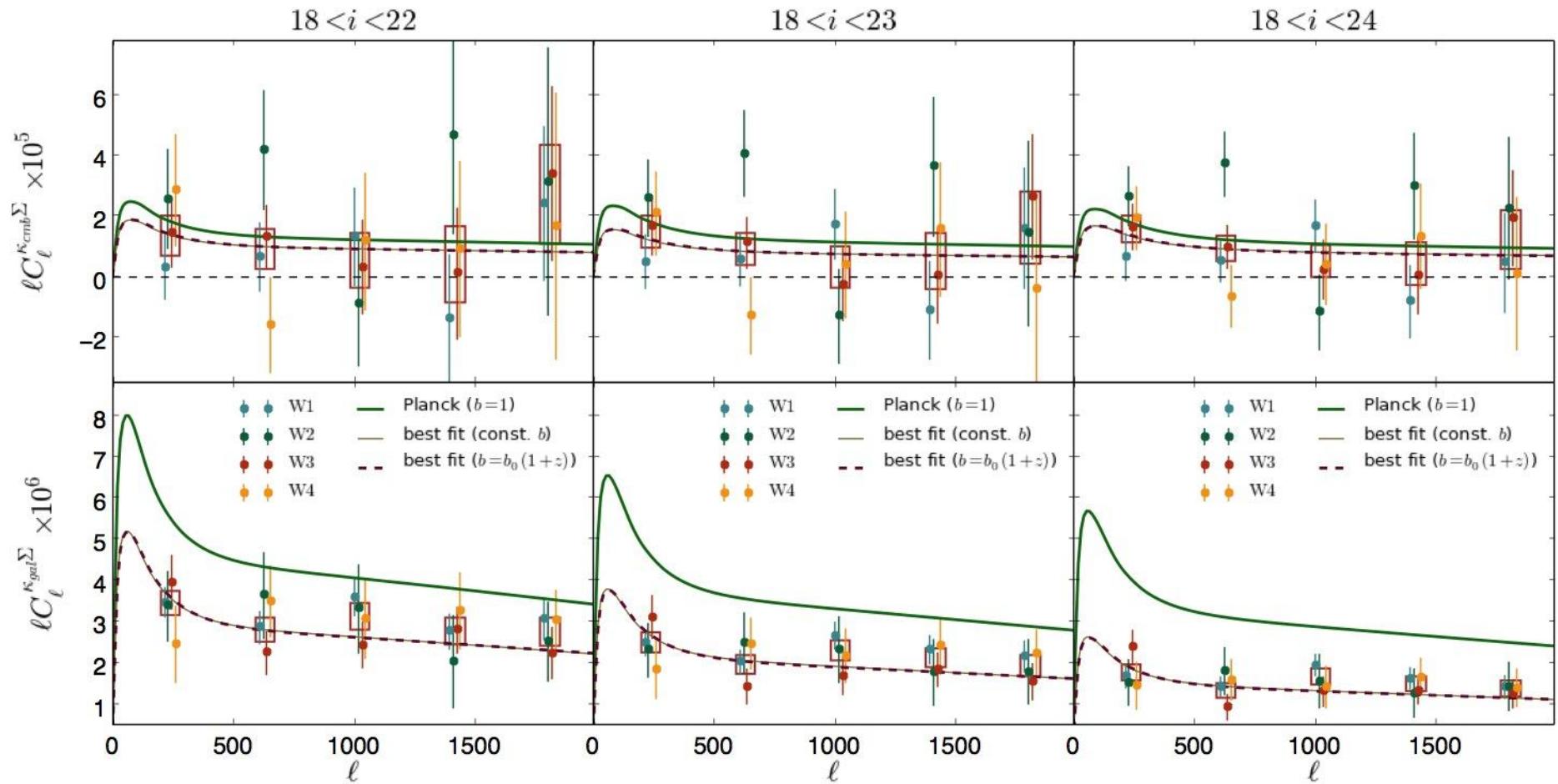
B $C_\ell^{\kappa_{\text{opt}} \Sigma} = m \frac{3}{2} \Omega_m H_0^2 \int d\eta b_\ell(\eta) W_f(\eta) \frac{g_{\text{opt}}(\eta)}{a(\eta)} P\left(\frac{\ell}{d_A}, \eta\right)$

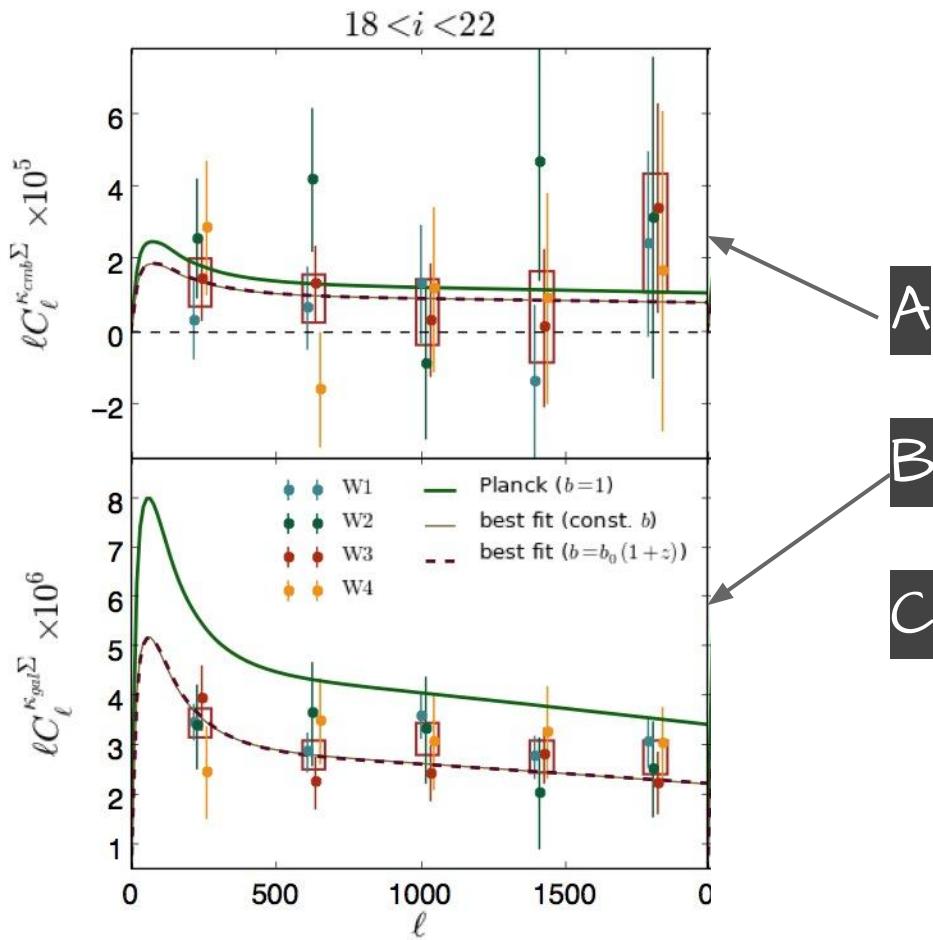
C $\frac{C_\ell^{\kappa_{\text{opt}} \Sigma}}{C_\ell^{\kappa_{\text{CMB}} \Sigma}} = m \frac{g_{\text{opt}}(\eta)}{g_{\text{CMB}}(\eta)}$

Formalism:
Vallinotto 2012
Das et al. 2013

Source Distribution & Lensing Kernels







A

$$C_\ell^{\kappa_{\text{CMB}} \Sigma} = \frac{3}{2} \Omega_m H_0^2 \int d\eta b_\ell(\eta) W_f(\eta) \frac{g_{\text{CMB}}(\eta)}{a(\eta)} P\left(\frac{\ell}{d_A}, \eta\right),$$

B

$$C_\ell^{\kappa_{\text{opt}} \Sigma} = \boxed{m} \frac{3}{2} \Omega_m H_0^2 \int d\eta b_\ell(\eta) W_f(\eta) \frac{g_{\text{opt}}(\eta)}{a(\eta)} P\left(\frac{\ell}{d_A}, \eta\right).$$

C

$$\frac{C_\ell^{\kappa_{\text{opt}} \Sigma}}{C_\ell^{\kappa_{\text{CMB}} \Sigma}} = \boxed{m} \frac{g_{\text{opt}}(\eta)}{g_{\text{CMB}}(\eta)}$$

Evidence of Multiplicative Bias

- ➡ A $2\text{--}4 \sigma$ evidence for the multiplicative bias ($m < 1$) in our deepest galaxy sample.
- ➡ Can potentially explain the disagreement between CFHTLenS shear 2-point function and Planck temperature measurements ($m \sim 0.9$ needed).
- ➡ Covariance dominated by the CMB lensing map noise at present.

JL, Ortiz-Vazquez, & Hill, 2016

